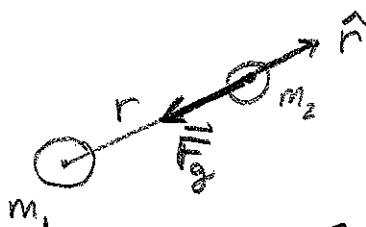


# LECTURE 33

- we worked out the inclined-plane torque problem and discussed bound vs. unbound orbits as seen from energy analysis. I'll let you get notes on the WA problem from a classmate if you missed it.

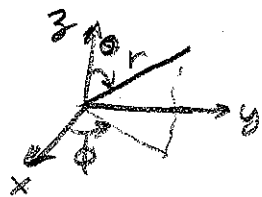
## Energy Analysis: (place $m_1$ at origin)

$$\vec{F}_g = -\frac{Gm_1m_2}{r^2} \hat{r}$$



In a good calculator III text (or course) you'll work out that in spherical coordinates

(physics spherical meaning of  $\theta$  and  $\phi$ )



$$\nabla g = \frac{\partial g}{\partial r} \hat{r} + \frac{1}{r} \frac{\partial g}{\partial \theta} \hat{\theta} + \frac{1}{r \sin \theta} \frac{\partial g}{\partial \phi} \hat{\phi}$$

switched from usual math conventions)

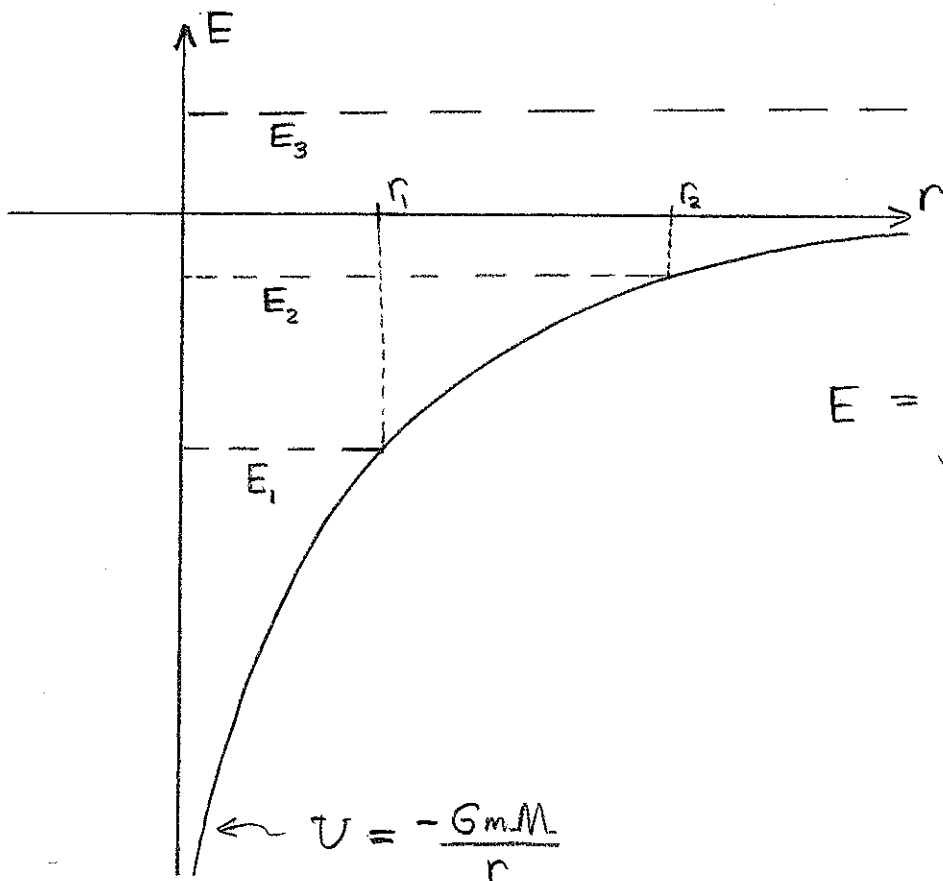
Comparison with  $\vec{F}_g$  leads us to conclude a Potential Energy function  $U$  with  $\vec{F}_g = -\nabla U$  has  $\frac{\partial U}{\partial \theta} = 0$  and  $\frac{\partial U}{\partial \phi} = 0$  so it only has dependence on the spherical radius  $r = \sqrt{x^2 + y^2 + z^2}$ .  
 Moreover, a few moments thought reveals,

$$U(r) = U_0 - \frac{Gm_1m_2}{r}$$

Often we set  $U_0 = 0$  which makes  $U(r \rightarrow \infty) = 0$ .



Continuing to analyze orbits



$$E = \frac{1}{2}mv^2 - \frac{GmM}{r}$$

energy for mass  $m$  subject to gravitational field of  $M$

E<sub>1</sub> The motion is bounded, since  $\frac{1}{2}mv^2 \geq 0$  we cannot allow  $r > r_1$ , since  $-\frac{GmM}{r} > E_1$  in that domain. It follows the motion cannot escape  $0 \leq r \leq r_1$ ,

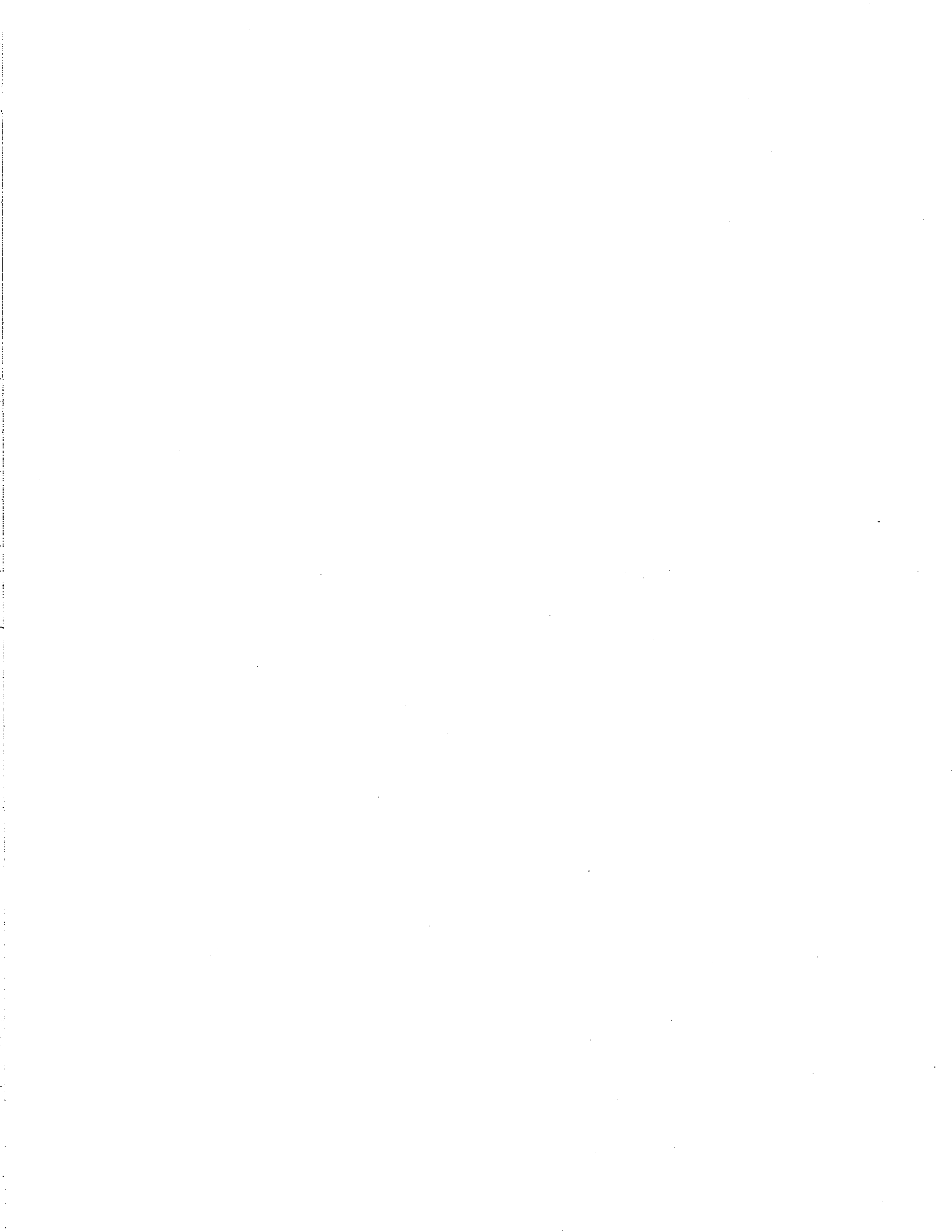
E<sub>2</sub> similar thinking, again bound between  $0 \leq r \leq r_2$  since  $KE > 0$ .

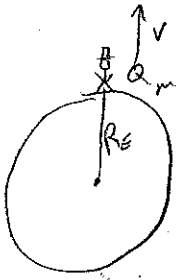
E<sub>3</sub> since total energy is positive we face no bound on  $r$ , as  $r \rightarrow \infty$  we find  $E \rightarrow \frac{1}{2}mv^2$ .

To complete the story we should think about angular momentum. Many planets travel essentially circular paths... in any event we need  $E = 0$  to escape so  $E = 0 = \frac{1}{2}mv^2 - \frac{GmM}{R_E} \rightarrow$

$$v_{\text{escape}} = \sqrt{\frac{2GM}{R_E}}$$

(for shooting straight up at Earth's surface)





$$\frac{1}{2} m v^2 - \frac{G M m}{R_E} - \frac{1}{2} m v_f^2 = 0$$

$$\sqrt{\frac{2 G M_E}{R_E}}$$

